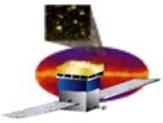


GLAST Large Area Telescope

LAT Science Working Group Review

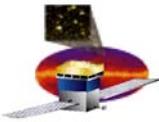
GRB Onboard Analysis

Jay Norris



GRB Onboard Analysis: Outline

- ❑ **Implementation Status — Progress on implementation of FSW algorithm for GRB trigger / localization**
- ❑ **Onboard performance**
- ❑ **Summary**



Implementation Status 1

□ Background

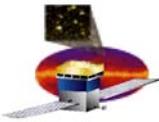
- ✓ Original IDL **trigger algorithm** — tested/refined over ~ 7 years: Utilizes the unbinned 2D spatial & 1D temporal coordinates and energy of detected event to generate probabilities for **2D+1D clusters of events**, as function of time.

An N-event “sliding window in time” is searched for significant 2D+1D event clusters.

- ✓ Upon significant trigger, IDL **localization algorithm** assembles all events between first & last significant cluster — that are mutually consistent with one position — in a T_{\max} window of ~ 150 s.

<< **Weight these photons by PSF(E) → Localization & Error** >>

- ✓ Notably, the mechanics of Onboard filters and knowledge of Onboard PSF(E) were not included.



Implementation Status 2

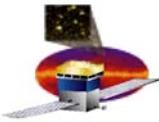
□ Onboard Simulations in C++

- ✓ IDL trigger & localization algorithms implemented into C++ code, with **extensive verification of C++ code** in Gleam environment.

Flavor of studies described here: **Onboard filters, Onboard PSF(E)**

- ✓ Specifically, the Onboard PSF(E) results from combining 2D track information to make 3D tracks. This additional “filter” reduces background rate ~ factor of 2 (375 Hz → 210 Hz).
- ✓ Refinement: In selection of best spatial cluster in a given “N-event sliding window”, **highest spatial probability cluster** rather than tightest cluster is chosen — increases cluster size.
- ✓ For ease of storage / retrieval: events in a **histogrammed map of the sky is maintained** over a relatively long window in time —

Allows possibility of accumulating GRB photons over longer than fixed ~ 10-s window, for localization algorithm.



Implementation Status 3

□ Conversion to FSW “C”

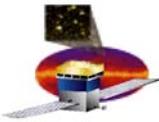
- ✓ Anticipated “C” code necessary to realize original algorithms into FSW is 80-90% implemented.

No problem foreseen in completing implementation.

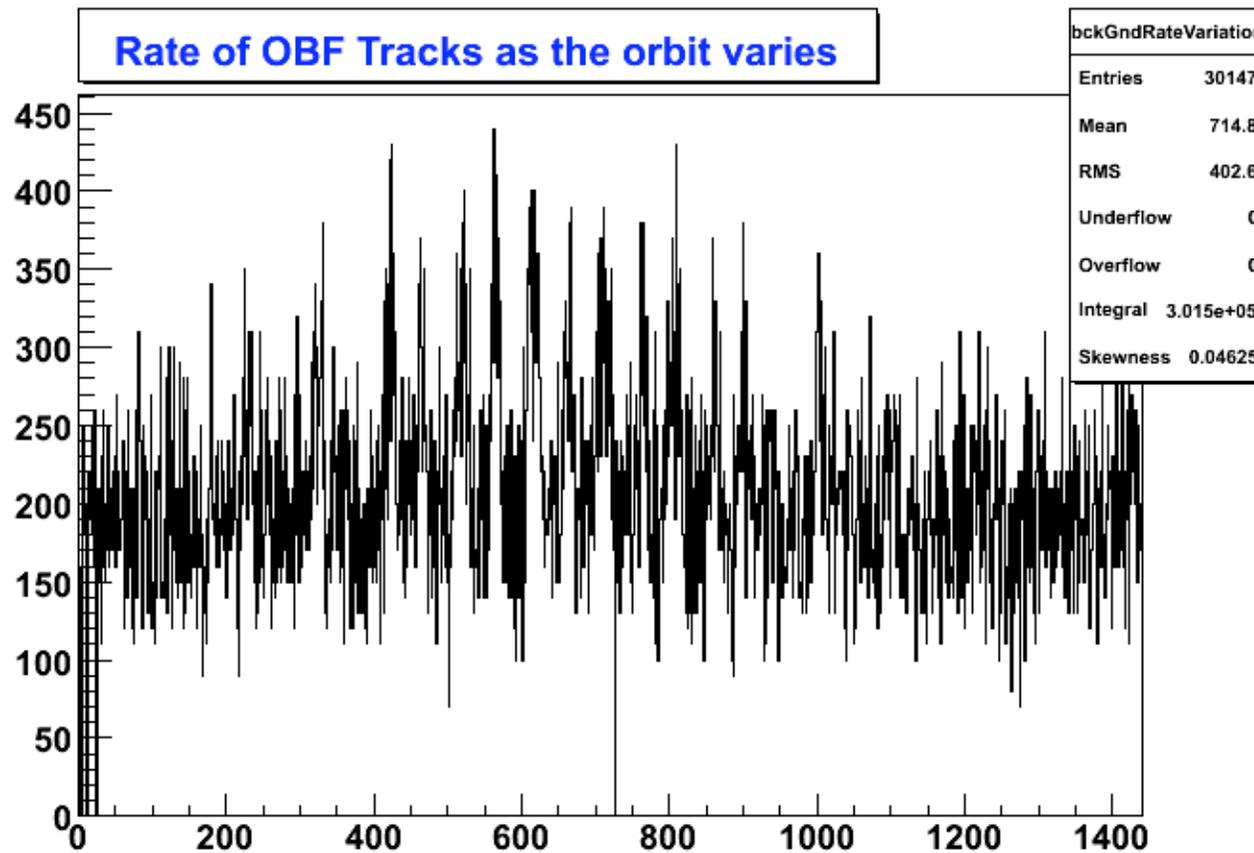
No “performance wall” for reasonable parameter choices.

✓ Computational intelligence:

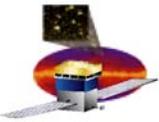
- Number of loops reduced — maintenance of prior computations: by adding effect of newest event, dropping oldest event.
Trig formulae allowed separation of Event{i} from Event{i+k}
<< Renders probability computation order N instead of N^2 >>
- “Modularized” temporal structures, reducing number of transcendental computations from 3 to 2 per event.
- Lists kept, storing transcendental computations that would have been done more than once.



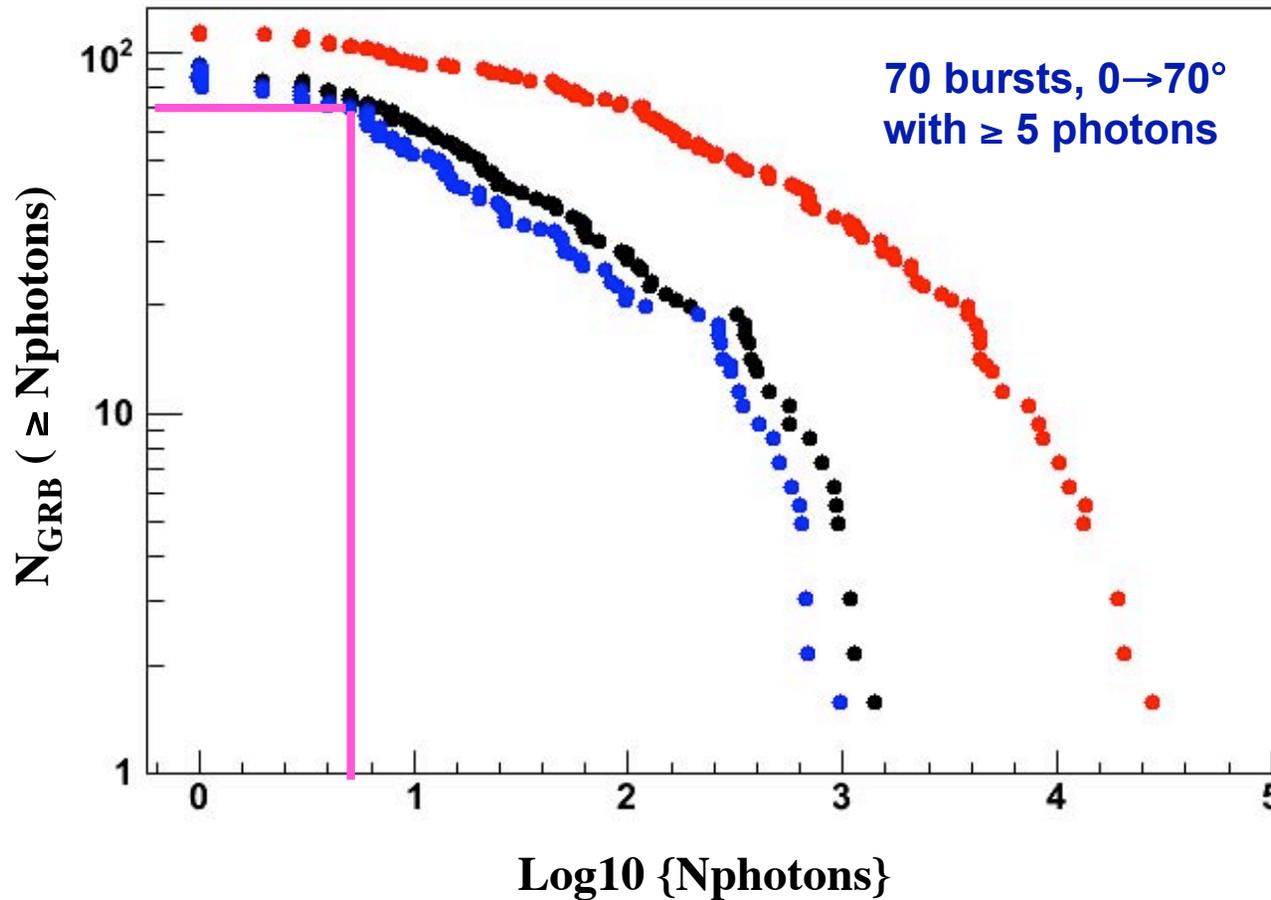
Background Rate for 3D Tracks



**Average upon requiring 3-D track ~ 210 Hz,
down from average “input” rate of ~ 375 Hz.**



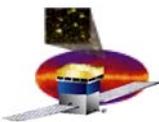
Sample with 3D tracks in LAT FOV



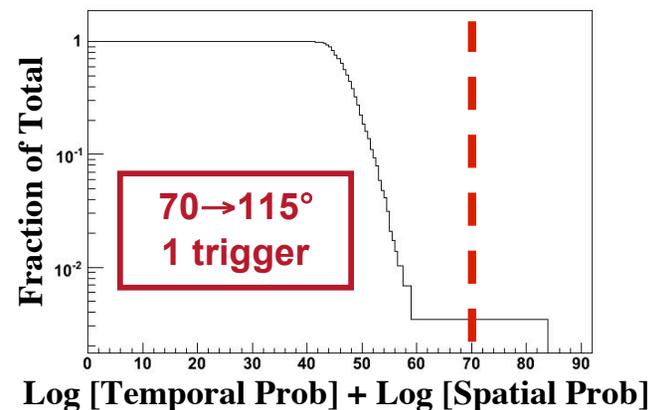
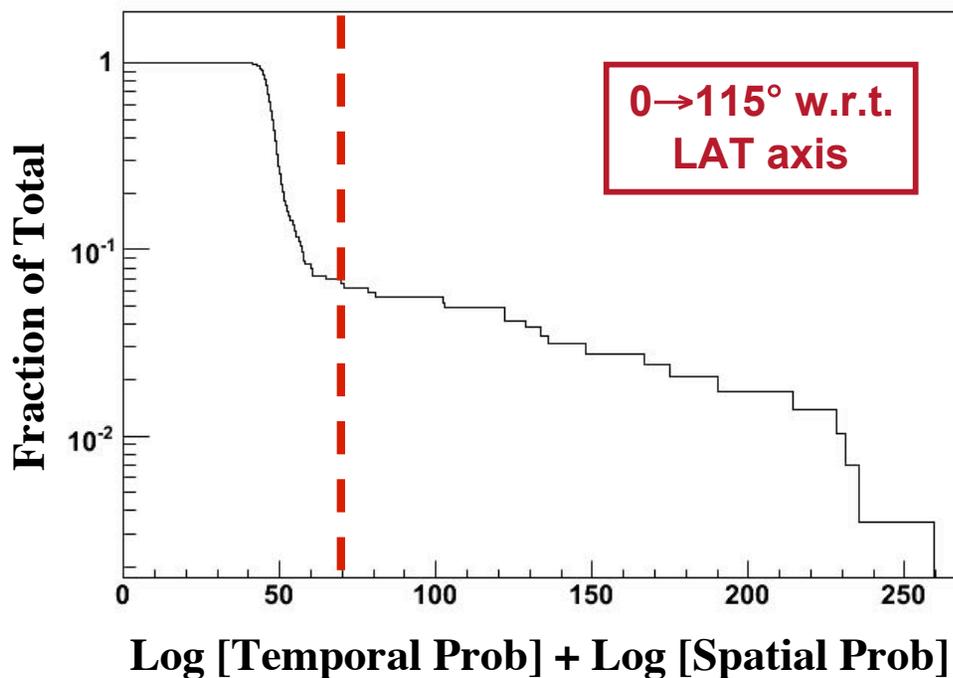
MC photons
pass OBF
have 3D track

Input: 290 bursts, isotropic
0→115° w.r.t. LAT axis

Measure trigger efficiency with:
Fixed background rate = 210 Hz,
Threshold ~ 1 false trigger/week

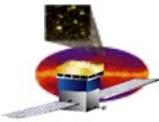


GRB Trigger Efficiency

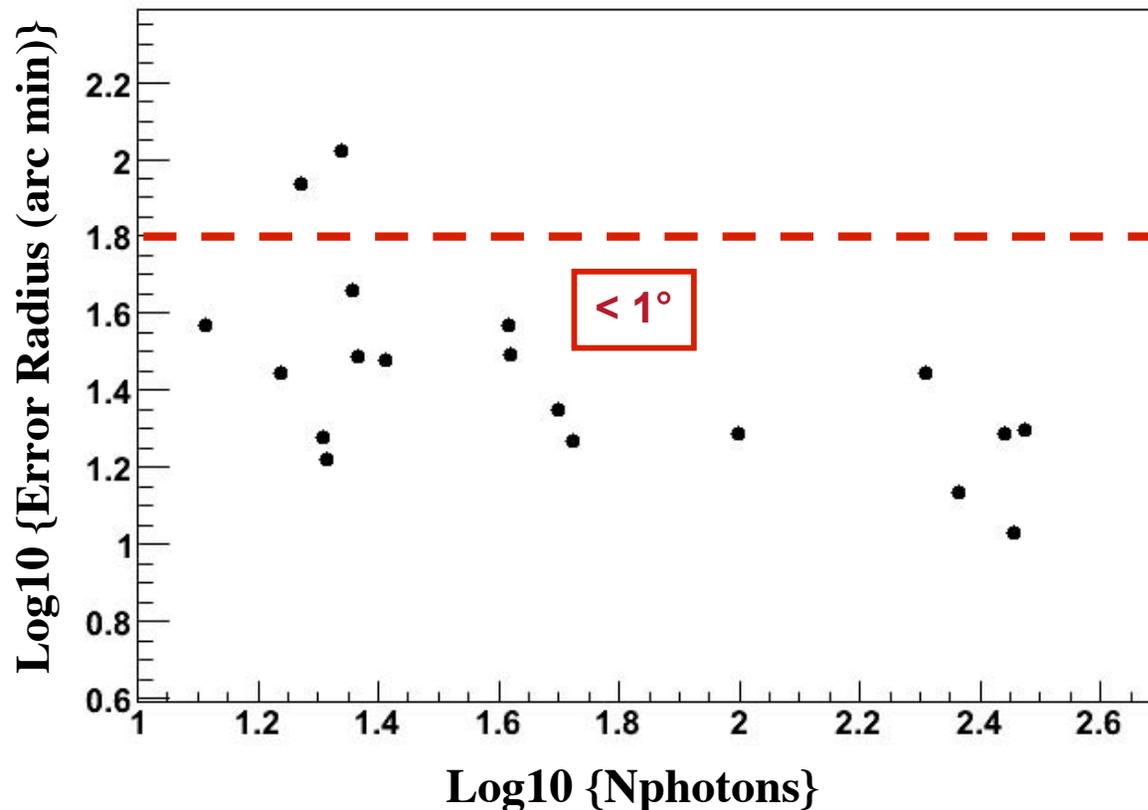


Roughly, Current Threshold:
LAT: $0.025 \text{ phot cm}^{-2} (> 30 \text{ MeV})$
(for 10-s localization list, E^{-2})
Corresponding BATSE threshold
 $F_p \sim 5 \text{ phot cm}^{-2} \text{ s}^{-1}$

- With $\text{Log}\{\text{Prob}\}$ Threshold = 70, efficiency ($0 \rightarrow 70^\circ$) is $\sim 16\%$ — 19 bursts trigger.
- Plateau on left hand side is due to background clusters.
- Fraction of GRBs with ≥ 5 detected photons in this sample is $\sim 25\%$ (~ 70 bursts).

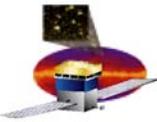


OnBoard Estimated Errors, Triggered Bursts

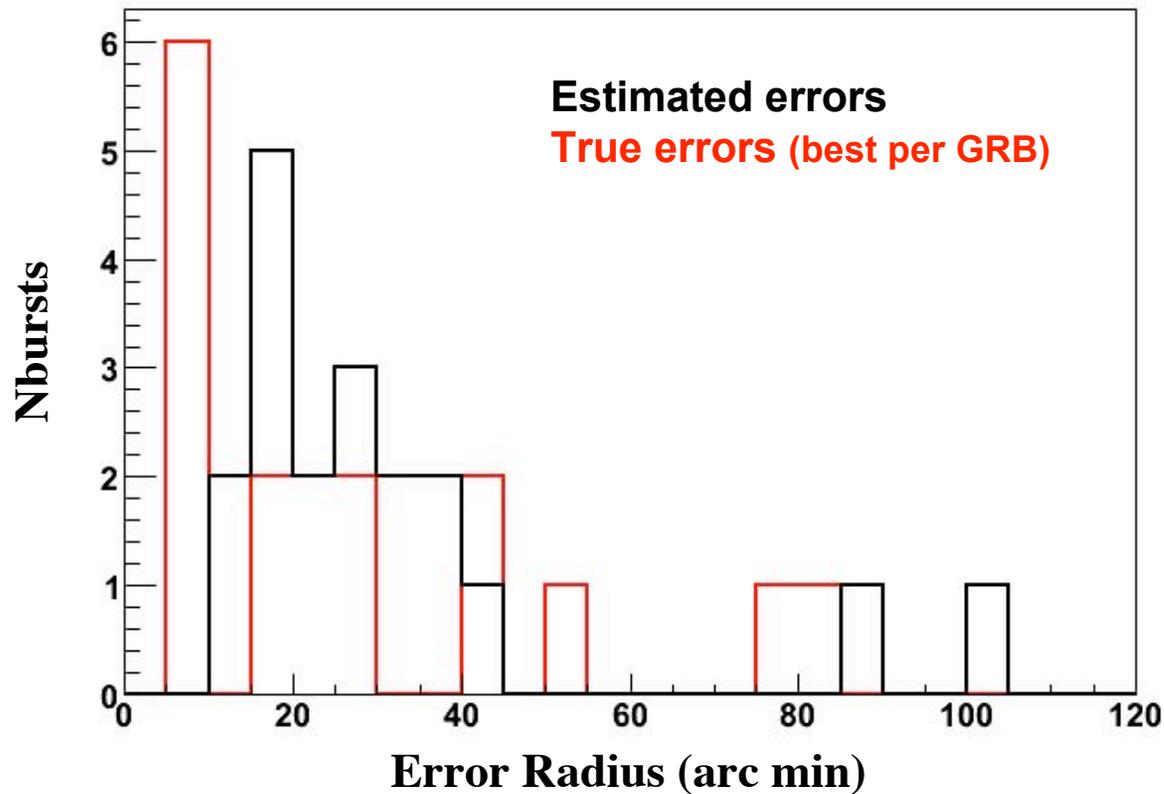


- Clusters formed from 40 most-recent events with 3D tracks
- Localization list < 10 s (but often, more photons available per burst)

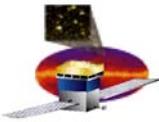
From 116-burst sample in 0→70° FOV:
~ 19 trigger, ~ 17 with Errors < 1°



Estimated vs. True Error, Triggered Bursts

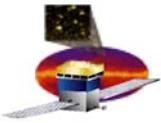


[True errors are often smaller than estimated, due to: non-Gaussian statistics of steep number spectra, and PSF(E)]



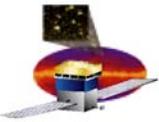
Summary: GRB Onboard Analysis

- Implementation has been rigorous reproduction of original algorithm over 9 months (IDL → C++ → “C”), with fixes and well-considered improvements.
 - FSW implementation is ~ 80-90% done. Several optimizations. Remainder is simple: We see no problem expeditiously completing algorithm.*
 - Currently demonstrated $\text{PSF}_{\text{OnBoard}} \sim 2 \times \text{PSF}_{\text{OnGround}}$, ~ independent of energy.
 - GRB yield ~ 16%, with threshold set for ~ 1 false trigger / week.
 - Current Onboard localizations: ~ 17 (per half year) with error radii $< 1^\circ$.
 - If there were a burst with 100 γ 's with E's > 1 GeV, Onboard error radius would be < 5 arc minutes — that meets the SRD. This is modern estimate, in line with current LAT performance characterization, and modified for Onboard PSF.
 - Expected Improvements. PSF: by improving quality of Onboard 3D tracks. Background Rate: Further study of filter veto bits (only to reduce background entertained by GRB algorithm).
- * Testing is acknowledged to be far and away the larger effort. Note: Testing in a hardware mockup is virtually independent of specific trigger and localization algorithm —which again, is nearly in place.

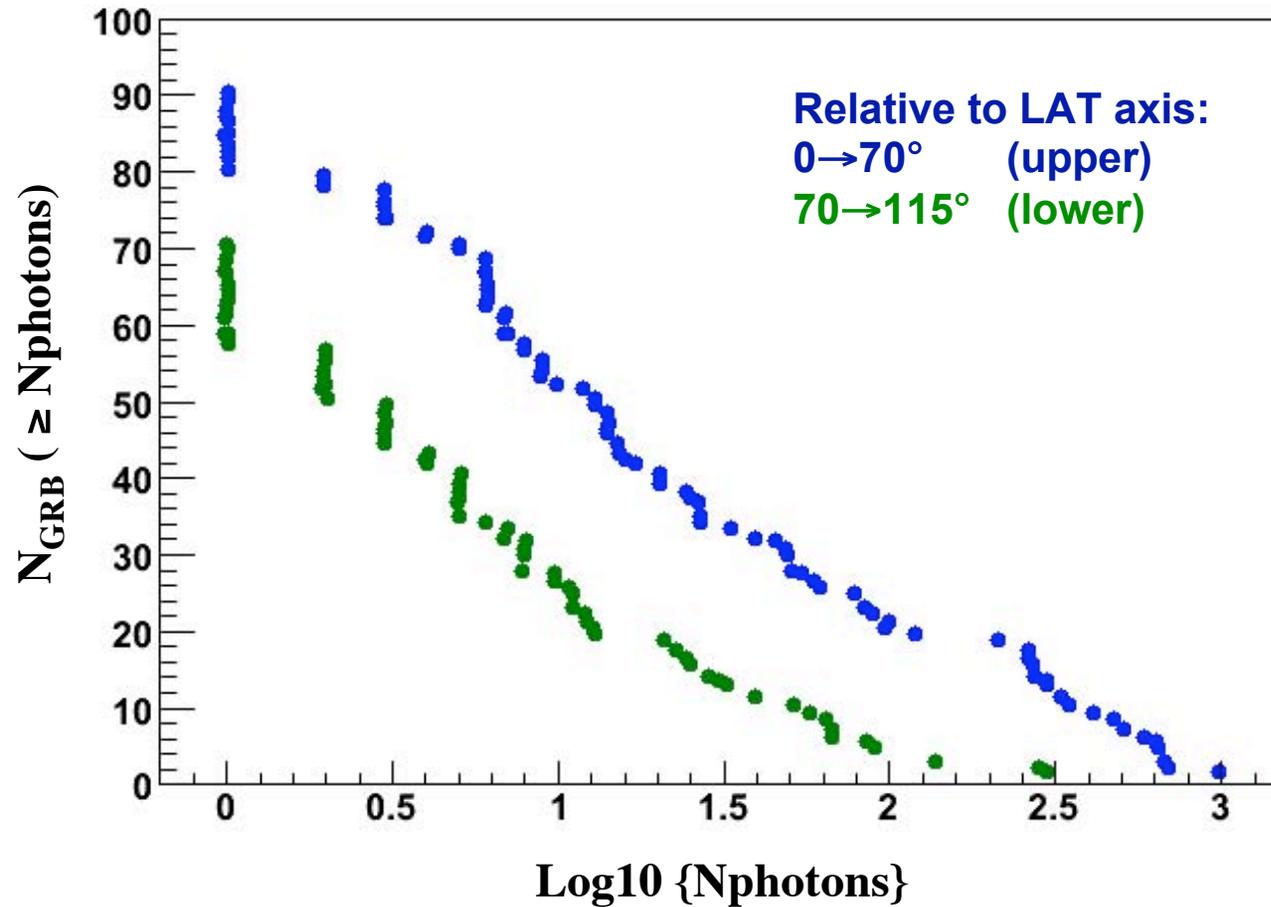


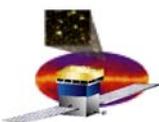
GRB Onboard Analysis

Backups

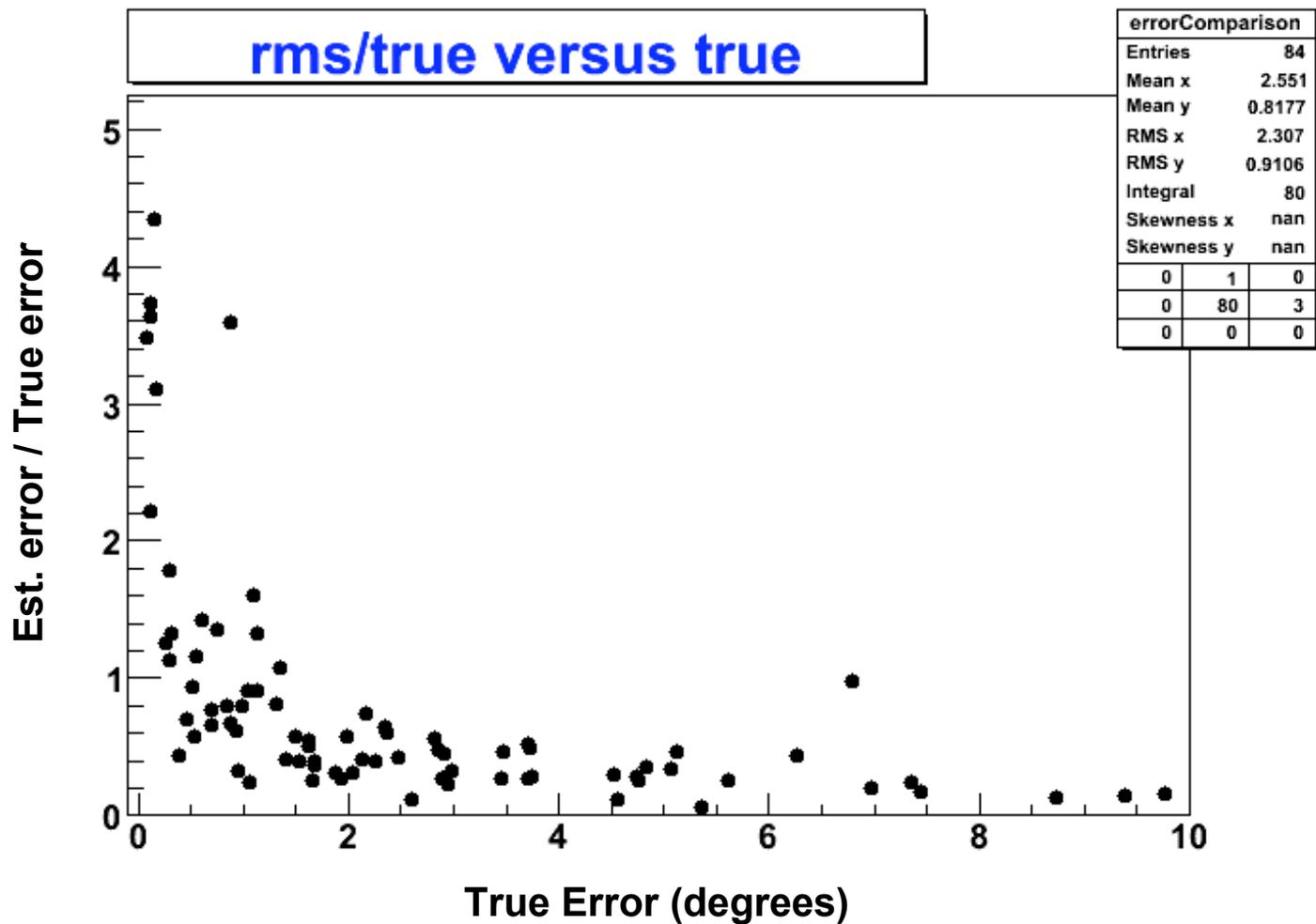


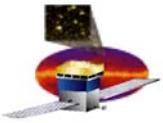
Sample with 3D tracks, 0→115°



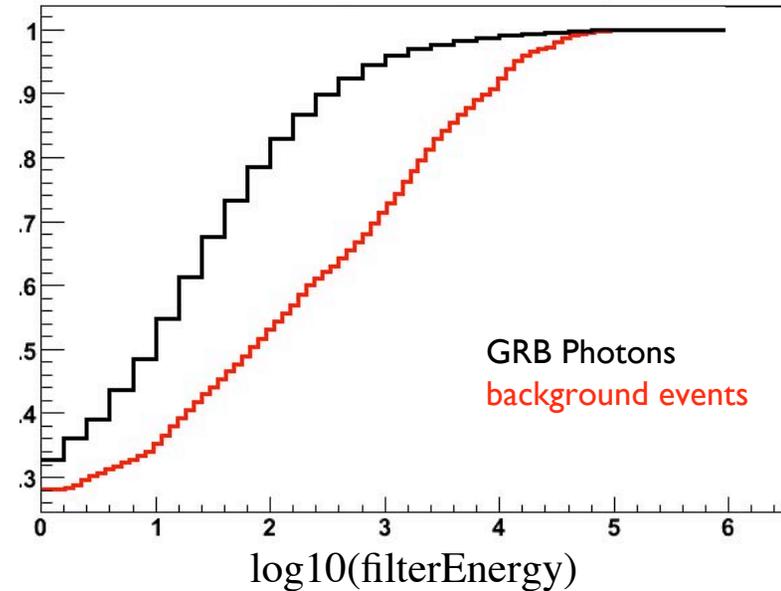
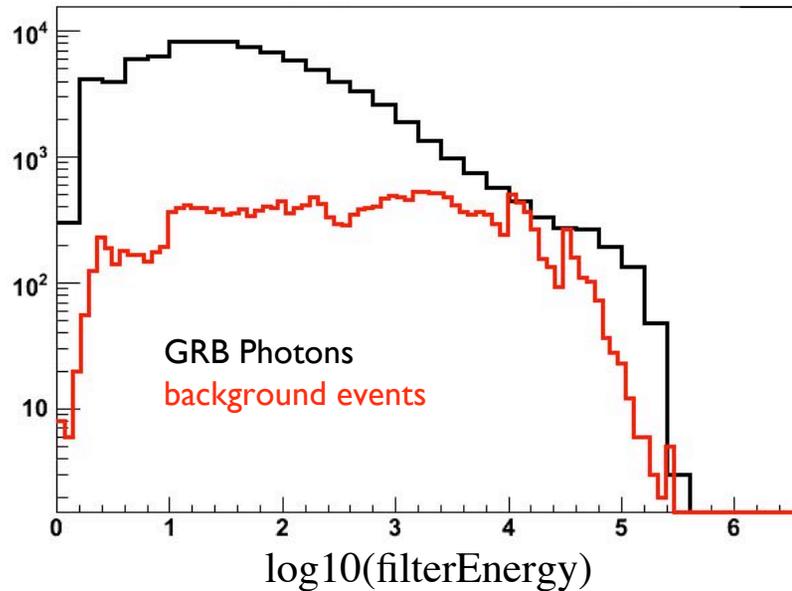


Estimated vs. True Errors

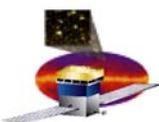




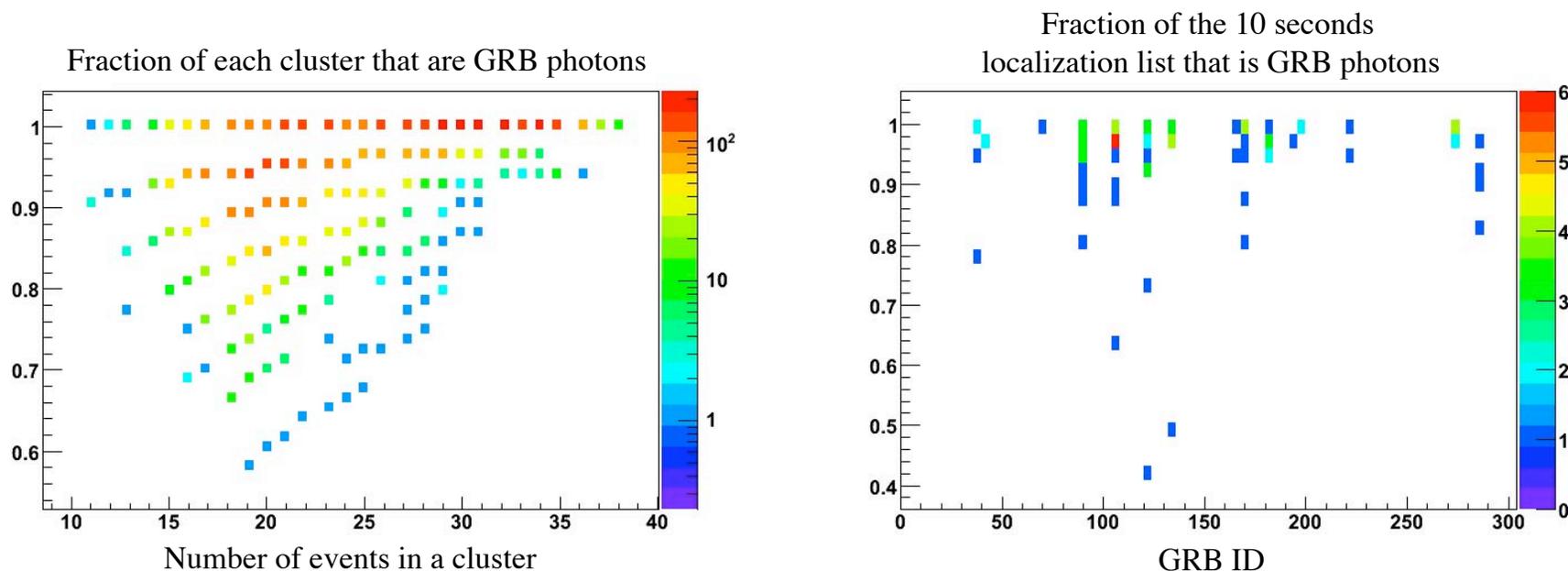
Filter Energy: Signal vs. Background



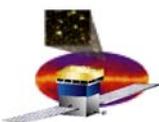
- Left (right) plot: differential (integral) E distributions of GRB γ 's and background.
- ~33% of GRB γ 's have zero filter energy (~28% of the background events).
- *If the PSF were energy-independent*, then with an energy cut we would always lose a larger fraction of GRB photons than background events.



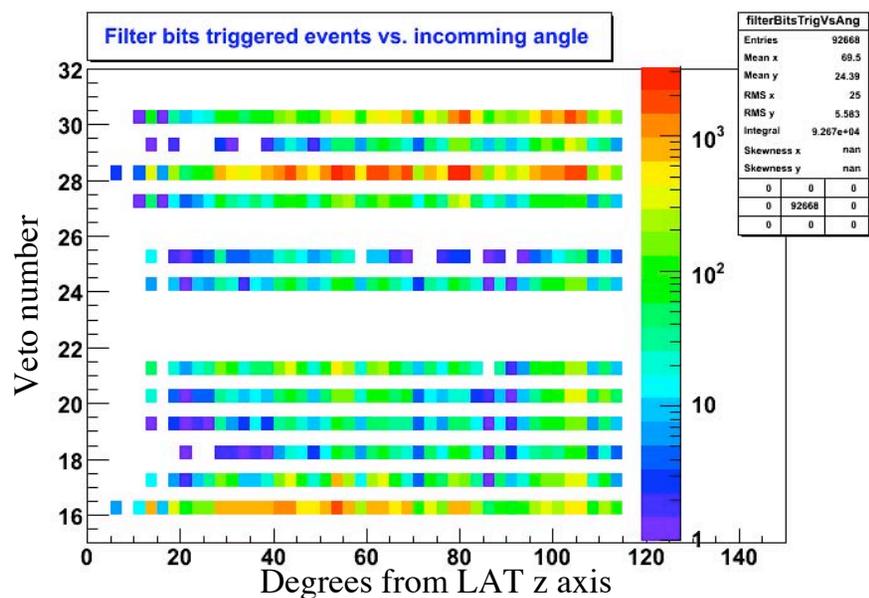
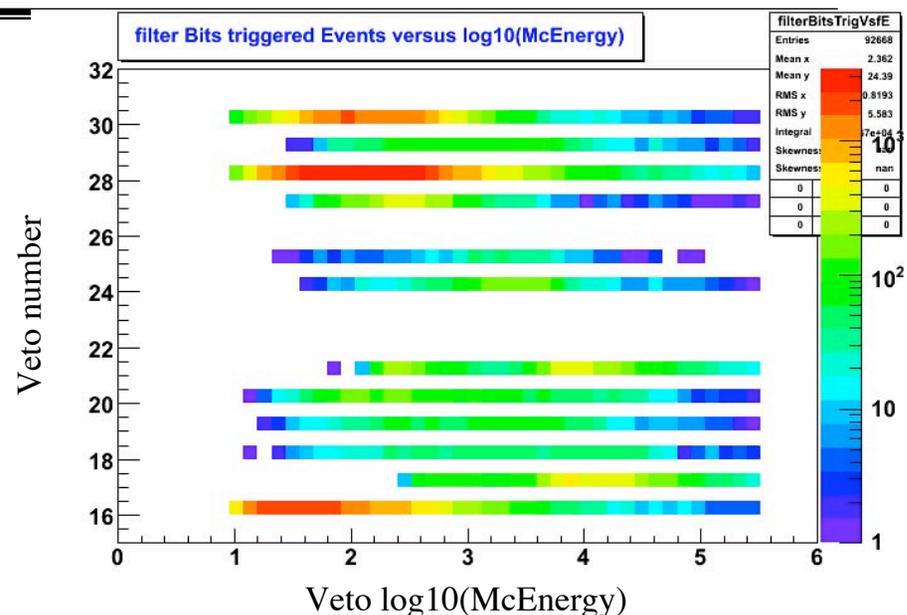
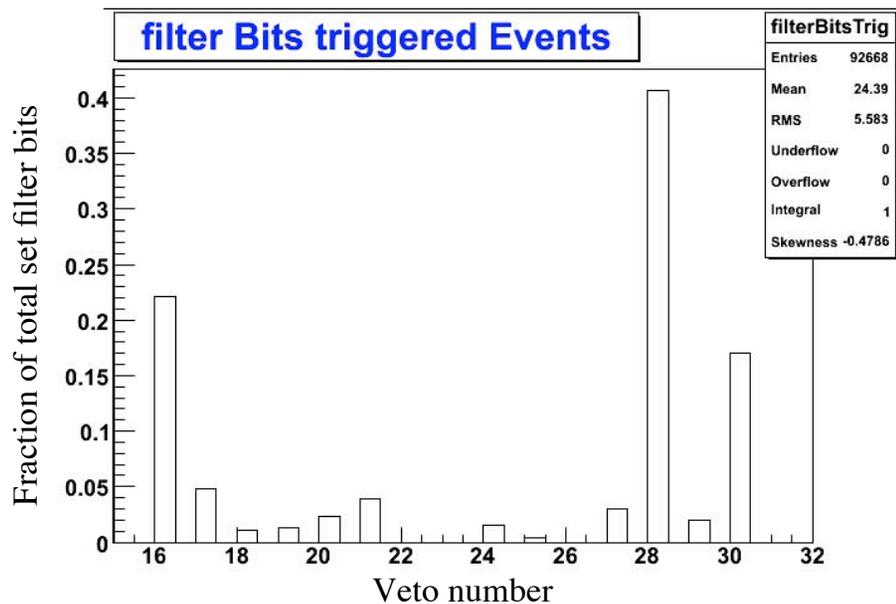
Purity of Clusters and Localization List



- All clusters have ≥ 10 events — no GRB trigger occurs with < 10 photons.
- Clusters are created from most-recent 40 events with OBF 3D tracks.
- Median **purity of Clusters (localization photon list)** is $\sim 95\%$ (96%) GRB photons.
- Threshold: 1 false trigger/week. Increasing threshold to one/35 days does not significantly change trigger efficiency (Log{Prob}: 70 \rightarrow 75).
- **Bottom Line: We have a relatively clean sample with the probability cut used.**



OnBoard Filter: Effects on GRB gammas



- Vetoes 28 and 30 are set (roughly) when the ACD is over threshold and there is no energy in the CAL
- Veto 16 is set when a projection passes through the space between the CAL and the ACD (the “skirt”)
- We are losing gammas that come from the side or below the LAT or are at low energies
- The 3D track filter is 77% efficient at retaining GRB gammas that trigger
- Vetos with no entries are currently turned off